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Bi-induced band gap reduction in epitaxial InSbBi alloys

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The properties of molecular beam epitaxy-grown InSb_{1-x}Bi_x alloys are investigated. Rutherford backscattering spectrometry shows that the Bi content increases from 0.6% for growth at 350 °C to 2.4% at 200 °C. X-ray diffraction indicates Bi-induced lattice dilation and suggests a zinc-blende InBi lattice parameter of 6.626 Å. Scanning electron microscopy reveals surface InSbBi nanostructures on the InSbBi films for the lowest growth temperatures, Bi droplets at intermediate temperatures, and smooth surfaces for the highest temperature. The room temperature optical absorption edge was found to change from 172 meV (7.2 μm) for InSb to ~88 meV (14.1 μm) for InSb_{0.976}Bi_{0.024}, a reduction of ~35 meV/%Bi. © 2014 Author(s). All article content, except where otherwise noted, is licensed under a Creative Commons Attribution 3.0 Unported License. [<http://dx.doi.org/10.1063/1.4902442>]

III-V-bismide alloys have attracted attention due to their interesting band structure, significant Bi-induced band gap reduction, and increase in spin-orbit splitting which make them suitable for optoelectronic devices operating in the near- and mid-infrared regions.¹⁻⁴ While III-SbBi alloys have been less extensively researched than III-AsBi alloys, interest in them is increasing, particularly for high power lasers and detectors in the important 2–5 and 8–14 μm ranges. There are only a few recent reports on the growth and properties of GaSbBi and related alloys, targeting the lower wavelength of these ranges.⁵⁻¹¹ Meanwhile, InSbBi was one of the first dilute bismides to be explored.¹²⁻¹⁶

In 1969, InSbBi alloys with a few percent Bi were identified by Jean-Louis and Hamon as a potentially important alternative to HgCdTe for long wavelength infrared sources and detectors operating in the 8–14 μm atmospheric transmission window.¹² Their pioneering work on InSbBi alloys involved investigations of bulk Czochralski-grown crystals. The lattice parameter was found to increase linearly with Bi content up to the maximum achieved of 2.4% of the anion sublattice.¹⁴ Optical absorption showed that the InSbBi single crystals exhibit respective band gap reduction of ~38 and ~30 meV/% Bi at 4 and 300 K.^{12,13} Rather surprisingly, these optical results have not been clearly reproduced for epitaxial material in spite of many reports on InSbBi over the last four decades.

In fact, reports on the band gap reduction for epitaxial InSbBi are rather scarce. They include a recent report of ~6 meV reduction for 0.4 atomic % Bi content; this low value was attributed to low substitutionality of Bi on group V lattice sites.⁵ In a report on metal organic vapor phase epitaxy (MOVPE)-grown InSbBi, promising indications of

band gap reduction were provided by photoresponse measurements, but experimental data on the composition were absent.^{17,18} The same study reported lattice contraction with respect to InSb, a result recently also obtained for InSbBi grown by liquid phase epitaxy (LPE) and molecular beam epitaxy (MBE).^{5,17,19} This contrasts with both the predictions of theoretical calculations²⁰ and also the lattice expansion found in the Czochralski-grown samples and other InSbBi alloy samples grown by MBE²¹ and MOVPE.²²

In this letter, we report the structural and optical properties of InSbBi alloys grown by MBE. The InSbBi epilayers were grown by solid-source MBE on 100 nm InSb buffer layers on undoped InSb(001) substrates. The Bi beam equivalent pressure flux was set to approximately 3.3×10^{-8} mbar using the beam monitoring ion gauge. The InSbBi epilayers were grown to a thickness of ~300–400 nm at a fixed growth rate of nominally 0.4 μm h⁻¹ using substrate temperatures ranging from 200 to 350 °C. Growth temperatures were determined using a pyrometer down to 240 °C. This enabled thermocouple readings to be calibrated and the calibration was extrapolated down to 200 °C.

The Bi content and lattice constant of the InSbBi films were determined using RBS with 3.72 MeV He²⁺ ions and HRXRD using a Philips X'Pert diffractometer equipped with a monochromatic Cu Kα₁ x-ray source (λ = 0.15406 nm). The surface morphology and composition of the InSbBi alloys was investigated by JEOL JSM-7001F FESEM with EDS and an Asylum research MFP-3D AFM in AC mode (tapping mode). The surface composition was also probed by hard x-ray photoemission spectroscopy (HAXPES) measurements. A photon energy, hν, of 4 keV was used and the instrumental broadening was 0.45 eV. Transmittance measurements were performed using a Bruker Vertex 70 V Fourier-transform infrared spectrometer, using a liquid nitrogen-cooled HgCdTe detector.

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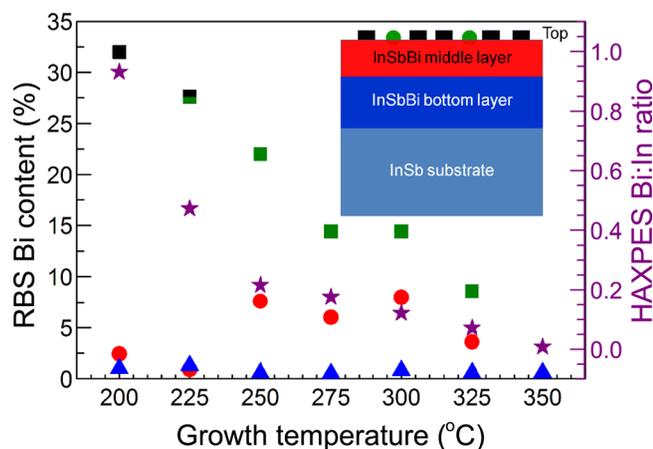


FIG. 3. The Bi content from RBS as a function of growth temperature in the “top,” “middle,” and “bottom” layers of the InSbBi epilayers represented, respectively, by squares, circles, and triangles. The inset shows a schematic diagram of the InSbBi films. The RBS Bi content is expressed as a proportion of the anion sublattice and uses the left axis. The Bi contents for the “top” layer InSbBi NS and Bi droplets should be taken as a qualitative indication of the trend of decreasing amount of Bi at the surface with increasing growth temperature rather than as quantitative values, as the RBS analysis assumes a planar compact film throughout. EDS indicates that the true Bi content of the NS is ~ 30 atomic % and for the Bi droplets is 100%. The Bi:In ratio from HAXPES is shown by the stars and uses the right axis.

levels. The curve-fitting, using Voigt functions, included the indium and bismuth oxide surface components, and their contribution was omitted from the Bi:In ratio. The Bi:In ratio was corrected to account for differences in orbital cross-sections and reflects the composition within the first 10–15 nm from the surface. For growth at 200 °C, the Bi:In ratio is close to 1 due to the presence of NS containing up to 40 atomic % Bi and regions of InSbBi films with Bi anion substitution of 2.4% and below. As the growth temperature is increased, the HAXPES Bi:In ratio decreases corresponding to greater Bi desorption and decreasing Bi droplet coverage. For the highest growth temperature of 350 °C, where NS and Bi droplets are absent the surface Bi:In ratio is consistent with the RBS-determined Bi content of the InSbBi film of 0.6%.

The Bi contents in the “middle” and “bottom” layers of the sample grown at 225 °C are, respectively, 0.9% and 1.3% (0.85% and 1.2% if only substitutional Bi is considered). The films grown at intermediate temperatures (250–325 °C) have thin (~ 60 –75 nm) “middle” layers with Bi content varying from 3.6% to 8.0% and ~ 300 nm-thick “bottom” layers with low Bi content of $\sim 0.5\%$ to 0.8%. Channeling RBS shows that all the samples are of high crystallinity with greater than 95% of Bi atoms on substitutional group V lattice sites in the bottom layers and also in the middle layers for growth at 200 and 225 °C, but only $\sim 25\%$ Bi substitutionality in the middle layers for growth between 250 and 325 °C.

The RBS evidence of a transition between the bottom layer and the higher Bi content middle layer for all but the sample grown at 350 °C is similar to previous observations for some GaSbBi samples grown at 250 °C.⁹ Similarly to previous discussions of this phenomenon in GaSbBi, it suggests that the growth mode changes after the growth of several tens of nanometers of InSbBi. The reason for this

change is currently unknown, but may be related to the requirement to accurately balance the growth temperature, V:III ratio and Bi flux during growth. The apparently spontaneous growth of a higher Bi content layer may be due to a real growth temperature change as a result of growth of a narrow band gap material on a wider band gap substrate while the apparent (thermocouple) temperature is unchanged. Alternatively Bi is known to have surfactant properties which could change the growth mode and Bi incorporation once sufficient Bi builds up on the surface.⁹

The HRXRD ω - 2θ scans of InSbBi samples grown at 200, 225, and 300 °C are shown in Fig. 4(a). The most intense peak is due to 004 reflections from the InSb substrate. The peak at lower Bragg angle corresponds to the InSbBi film and indicates that, under these growth conditions, the lattice expands when Bi is incorporated, in contrast to some earlier reports on MBE grown InSbBi and GaSbBi alloys.^{6,19} The diffraction peak corresponding to the InSbBi epilayer shifts towards lower angle with decreasing growth temperature and increasing RBS Bi content. In order to study the strain state of the epilayers, asymmetric 115 reciprocal space maps (RSM) were performed. Fig. 4(b) shows the HRXRD RSM around the 115 reflections of the InSb substrate and the InSbBi film grown at 200 °C. Alignment of the InSb and InSbBi peaks along Q_x indicates that the film is fully strained to the InSb in-plane lattice constant. Asymmetric maps indicate that the InSb_{1-x}Bi_x films are all completely strained.

The XRD results are consistent with the RBS results. For the sample grown at 200 °C, the 004 peak at $\sim 28.36^\circ$ corresponds to the layer with 2.4% Bi and the intensity between that peak and the substrate peak corresponds to the layer containing 1.0% Bi. Intensity associated with these two layers is also apparent in the RSM in Fig. 4(b). The XRD peak for the sample grown at 225 °C corresponds to a Bi content of $\sim 1.0\%$ Bi, but is broader than expected for a single layer, consistent with the RBS evidence of two separate layers containing 0.85% and 1.2% Bi. For the samples grown at higher temperatures (250–325 °C), the XRD data have a shoulder on the low angle side of the substrate peak, the width of which correlates with the Bi content indicated by RBS for the bottom layer and for the single layer with uniform Bi content for growth at 350 °C. No clear XRD signals were obtained from the higher Bi content layers, consistent with the low Bi substitutionality, and lower crystallinity indicated by RBS for the middle layers.

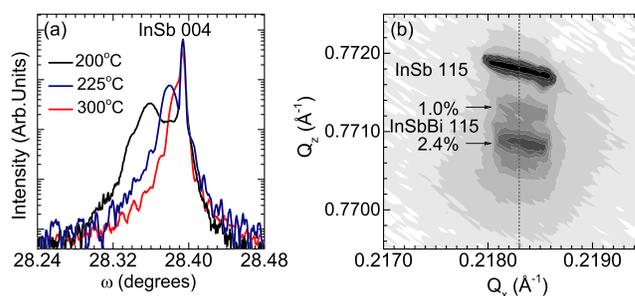


FIG. 4. HRXRD scans of InSbBi films on InSb substrates. (a) ω - 2θ scans of InSbBi samples grown at temperatures of 200, 225, and 300 °C. (b) Reciprocal space map around the 115 reflection of the InSbBi sample grown at 200 °C.

The InSbBi lattice constant was calculated by considering the tetragonal distortion of the InSbBi films, using the Poisson ratio and elastic constants of InSb.^{26,27} The lattice parameters in the inset of Fig. 2(b) show a linear trend in accordance with Vegard's law. If this linear trend is extrapolated to 100% Bi, the lattice parameter corresponding to zinc-blende InBi is $6.626 \pm 0.010 \text{ \AA}$. This is smaller than the 7.292 \AA (Ref. 28) and 7.024 \AA (Ref. 29) from extrapolating InPBi and InAsBi lattice parameters, respectively, and smaller than the theoretically predicted zinc-blende InBi lattice constant (at 0K) of 6.686 \AA (Ref. 20), but in agreement with the value extrapolated from bulk grown InSbBi samples of $6.638 \pm 0.011 \text{ \AA}$ (from our linear fitting of the data in Ref. 14).

The optical properties of the InSbBi samples as a function of Bi content were studied using transmittance measurements. The transmission data from each sample were divided by the transmission from an InSb substrate so that the remaining signal corresponds to the InSbBi epilayer. The derived absorption spectra are shown in Fig. 5. The inset of Fig. 5 shows the measured absorption edge energy as a function of Bi content. The optical band gap red shifts with increasing Bi content. The sample grown at 200°C shows two absorption edges corresponding to the layers with Bi content of 1.0% and 2.4%. The absorption between the two extrapolations can be accounted for by the $\sim 60 \text{ nm}$ of the film in which the Bi content changes from 1.0% to 2.4%. When the Bi content is increased to $x=0.024$, the absorption edge decreases in energy to $\sim 88 \text{ meV}$, corresponding to a wavelength of $14.1 \mu\text{m}$. The dashed line in the inset of Fig. 5 shows the band gap reduction corresponding to the virtual crystal approximation (VCA). The VCA band gap reduction was calculated using the "negative band gap" of semi-metallic zinc-blende InBi of -1.7 eV from density functional theory,²⁰ which is $19.1 \text{ meV}/\% \text{Bi}$ consisting of $18.8 \text{ meV}/\% \text{Bi}$ lowering of the conduction band minimum (CBM) and $0.3 \text{ meV}/\% \text{Bi}$ increase of the valence band maximum (VBM). The

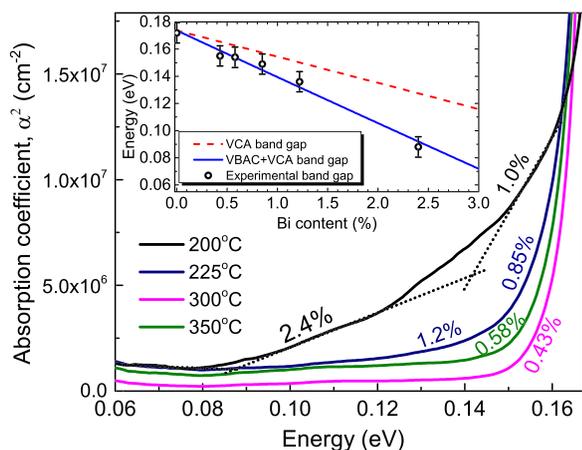


FIG. 5. The absorption spectra for $\text{InSb}_{1-x}\text{Bi}_x$ films. The composition of each film is given in the figure. Where the film comprises multiple layers, these may be associated with different absorption edges, as indicated. The inset shows the optical band gap versus Bi content determined from the absorption spectra (open circle). The band gap of $\text{InSb}_{1-x}\text{Bi}_x$ is calculated assuming the VCA variation of both the CBM and the VBM (dashed line) and by assuming the VCA variation of the CBM and VBAC between the valence bands and the Bi impurity level (solid line).

experimentally observed band gap reduction ($35 \text{ meV}/\% \text{Bi}$) exceeds the theoretical prediction of the VCA.

In order to model the experimental data, the valence band anticrossing (VBAC) approach was used. The observed band gap is well reproduced by the $12 \times 12 \mathbf{k}\cdot\mathbf{p}$ model described by Alberi *et al.* for GaAsSb and GaAsBi.³⁰ The VBAC interaction is included between localized Bi $6p$ -like states and the host valence bands. The Bi localized state energy level used is 1.2 eV below the InSb VBM. This value was chosen by considering the reported Bi level of 0.4 eV below the VBM in GaAs³⁰ and that the VBM of GaAs is 0.8 eV below that of InSb,²⁶ assuming that the Bi level is approximately equal in different materials on an absolute energy scale. Using this Bi level energy, the interaction strength required to fit the data is 1.35 eV . If the Bi level was closer to the VBM (for example, a value of 0.18 eV below the VBM of GaAs was calculated using tight binding in Ref. 31), a correspondingly lower value of interaction strength would be required. According to this VBAC modeling, a band gap reduction of $33.9 \text{ meV}/\% \text{Bi}$ is made up of $18.8 \text{ meV}/\% \text{Bi}$ VCA lowering of the CBM and $15.1 \text{ meV}/\% \text{Bi}$ due to the upward shift of the VBM predominantly due to the VBAC interaction. This amount of band gap reduction per $\% \text{Bi}$ for MBE grown material is very similar to the $30 \text{ meV}/\% \text{Bi}$ previously reported for Czochralski-grown bulk InSbBi alloys.¹³

In conclusion, $\text{InSb}_{1-x}\text{Bi}_x$ thin films have been grown by MBE. The Bi incorporation induces lattice dilation, in contrast to some of the previous reports. Extrapolation of the InSbBi lattice parameter to 100% Bi suggests a zinc-blende InBi lattice parameter of 6.626 \AA . The optical absorption edge decreases at a rate of $\sim 35 \text{ meV}/\% \text{Bi}$ from 172 meV ($7.2 \mu\text{m}$) for InSb to $\sim 88 \text{ meV}$ ($14.1 \mu\text{m}$) for $\text{InSb}_{0.976}\text{Bi}_{0.024}$. This makes InSbBi alloys suitable for sources and detectors operating in the $8\text{--}14 \mu\text{m}$ atmospheric transmission window. Further work is required to achieve InSbBi films with uniform Bi content up to 2.4% by exploring the effects of other MBE growth parameters.

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